# Advanced Image Combine Techniques

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#### Important Equation 1 of 22 (Joke! ☺)

#### Standard deviation of a discrete random variable or data set

The standard deviation of a discrete random variable is the root-mean-square (RMS) deviation of its values from the mean.

If the random variable X takes on N values  $x_1, \dots, x_N$  (which are real numbers) with equal probability, then its standard deviation  $\sigma$  can be calculated as follows:

- Find the mean, x̄, of the values.
- 2. For each value  $x_i$  calculate its deviation  $(x_i \overline{x})$  from the mean.
- 3. Calculate the squares of these deviations.
- Find the mean of the squared deviations. This quantity is the variance σ<sup>2</sup>.
- Take the square root of the variance.

This calculation is described by the following formula:

$$\sigma = \sqrt{\frac{1}{N} \sum_{i=1}^{N} (x_i - \overline{x})^2},$$

where  $\overline{x}$  is the arithmetic mean of the values  $x_i$ , defined as:

$$\overline{x} = \frac{x_1 + x_2 + \dots + x_N}{N} = \frac{1}{N} \sum_{i=1}^{N} x_i.$$

If not all values have equal probability, but the probability of value  $x_i$  equals  $p_i$ , the standard deviation can be computed by:

$$\sigma = \sqrt{\frac{\sum_{i=1}^N p_i(x_i-\overline{x})^2}{\sum_{i=1}^N p_i}}, \text{and}$$
 
$$s = \sqrt{\frac{N'\sum_{i=1}^N p_i(x_i-\overline{x})^2}{(N'-1)\sum_{i=1}^N p_i}},$$

where

$$\overline{x} = \frac{\sum_{i=1}^{N} p_i x_i}{\sum_{i=1}^{N} p_i},$$

and N' is the number of non-zero weight elements.

The standard deviation of a data set is the same as that of a discrete random variable that can assume precisely the values from the data set, where the point mass for each value is proportional to its multiplicity in the data set.

#### Outline

- Some History
- The Imager's "Enemies" artifacts!
- The Basic Raw Image Types
- Image Calibration Workflow
- Common Combine Algorithm Choices
- Combine Algorithm Details
- Choosing the Right Combine Algorithm
- Some New Algorithm Ideas

## History

- Film to CCD transition started in the 90's
- CCDs produced better results with shorter and fewer images
- Stacking CCD images was quickly adopted
- New software (DDP) was invented to show the entire dynamic range
- Result: Unprecedented faint detail BUT with artifacts

#### The "Enemies"

- Cosmic Ray Hits
- Hot/Cold Pixels
- Bad Columns
- Satellite Trails
- Plane Trails
- Asteroids
- Dust Motes, etc.

#### Some Basics

- Four Raw Image Types
  - Bias Frames
  - Dark Frames
  - Flat Field Frames
  - Light Frames
- Constraints
  - Temperature
  - Time

#### Bias Frames

- Consists of:
  - Bias Level can vary with time
  - Bias structure may change slowly over time
  - Read-out Noise

#### Dark Frames

- Consists of:
  - Bias level
  - Bias structure
  - Read-out noise
  - Dark-current



#### Flat-Field Frames

- Light frames used to correct variations in:
  - Illumination
  - Pixel to Pixel Sensitivity
- Common ways to create flat-field frames:
  - Light Box
  - Flat target
  - Twilight exposures

## Light Frames

- Consist of:
  - Light collected from imaging target/sky
  - Dark Current
  - Bias level, bias structure, and readout noise.

## Quick Data Collection Tips

- 1. Take all frames at the same Temperature.
- 2. Dark frame count = Light Frame Count
- 3. Dark Frame duration = Light Frame Duration
- 4. "Dither" all Light Frames and Flat Frames
- 5. Flat Frames should be at/near Focus
- 6. Flat Frames should be at Light Frame Camera Orientation
- 7. Flat Frames for every filter!

## Image Calibration Workflow

- Combine Bias Frames to Create a Master Bias
- 2. Combine Dark Frames to Create Master Dark
- 3. Subtract Master Bias from Each Flat Frame
- 4. Normalize then Combine Flat Frames
- 5. Subtract Master Dark from Light Frames
- 6. Apply Normalized Flat Field to Light Frames
- 7. Align Light Frames
- 8. Combine Light Frames

# Common Combining Algorithms

- Average
- Median
- Min/Max Clip
- Sigma Clip
- SDM (Standard Deviation Masking)

## Average

- Algorithm:
  - Sum of pixels divided by count of frames
- Advantage:
  - Highest Signal/Noise
- Disadvantage:
  - Artifacts are not rejected
- Example:
  - -5,7,12 = (5+7+12)/3 = 24/3 = 8

#### Median

- Algorithm:
  - Sort pixels from highest to lowest value and pick the one in the middle.
- Advantage:
  - Best noise rejection
- Disadvantage:
  - S/N is lower than Average combine.
- Example:
  - Median of 10,12,25,35,54 is 25

## Min/Max Clip

- Algorithm
  - Throw away the highest and lowest pixel values and average the rest
- Advantage
  - Rejects most artifacts
- Disadvantages
  - Still can leave some artifacts
  - Works best with 6+ images.
- Example
  - 10,12,14,18,20,30 Min/Max Clip = (12+14+18+20)/4 = 16

# Sigma Clip

- Algorithm:
  - Calculates the average and standard deviation of the pixels. Averages pixels that are within a Sigma factor range of the average.
- Advantage:
  - Strong noise rejection
- Disadvantages:
  - Needs many images (10+) to work best
  - Bad or failed rejections can occur.

# Sigma Clip Example

Suppose the pixel set consists of the values 3, 7, 7, and 19

Step 1: Calculate mean:

$$(3+7+7+19)/4=9$$

Step 2: Calculate deviation from mean:

$$3 - 9 = -6$$

$$7 - 9 = -2$$

$$7 - 9 = -2$$

$$19 - 9 = 10$$

Step 3: Square each deviation

$$(-6)*(-6) = 36,$$

$$(-2)*(-2) = 4$$

$$(-2)*(-2) = 4$$

$$(10)*(10) = 100$$

Step 4: Find mean of all deviations

$$(36 + 4 + 4 + 100) / 4 = 36$$

Step 5: Calculate Standard Deviation by

taking the square root:

$$SQR(36) = 6$$

Step 6: S is the user definable Sigma

factor. Let S=0.5 in this case

Step 7: Reject values outside of Mean – S

\* Standard Deviation:

$$Min = 9 - (0.5 * 6) = 6$$

$$Max = 9 + (0.5 * 6) = 12$$

In values: 3, **7**, **7**, and 19:

Reject 3 and 19.

Step 8: Average the remaining values

$$(7 + 7) / 2 = 7$$

## SDM Example

Suppose we have a stack of 3x3 pixel images

Step 1: calculate the mean, median and standard deviation values:

4	3	1	3	4	5		3	3	1
4	7	8	6	4	4		4	4	5
7	2	1	2	1	2		13	2	1
	Mean		N	Median			Standard Deviation		

Step 2: calculate the average of the standard deviations: (3+3+1+4+4+5+13+2+1)/9 = 36/9 = 4

### Recomendations

Action	Combine Method					
Create Master Bias or Master Dark	SDM: Factor=1.0, Passes=1-3, Normalization=None, Ignore Black Pixels=No, Despeckle=NO					
Create Master Flat	1) Median after Normalization 2) SDM: Factor=1.0, Passes=1-3, Normalization=Linear, Area=30%, Ignore Black Pixels=No, Despeckle=Yes					
Combine Light Frames	1) SDM: Factor=1.0, Passes=3, Normalization=Linear, Area=30%, Ignore Pixels over 1/3 well capacity, Ignore Black Pixels=Yes, Despeckle=Yes (Sigma factor 1.0)					

#### SDM

#### Algorithm:

- Creates median and average combined images
- Creates a standard deviation image (mask)
- Based on the standard deviation at an X,Y pixel either the median or the average pixel is used.
- Details: <u>http://www.gralak.com/Sigma/SigmaClippingAlgorithmDesignSummary.pdf</u>

#### Advantages:

- Works well with as little as 3 images, very strong noise rejection
- Not "fooled" by random sets of bad pixels

#### Disadvantages

Requires more resources (CPU + memory)

## Improved SDM

- Mix SDM with other algorithms (e.g. Min/Max Clip, then SDM)
- Standard deviation values vary by level
- Better normalization algorithm (use median of block regions)
- Automatic rejection of pixels surrounding a cluster of bad pixels in a particular image

## Summary and Conclusion

- The imager's "Enemies"
- The basic raw image types
- Image calibration workflow
- Common combine algorithm choices
- Combine algorithm details
- Choosing the right combine algorithm
- Some new algorithm ideas